

A Scalable AMR Gravity Solver for ENZO-E (Extreme-scale ENZO)

Michael L Norman and James Bordner

San Diego Supercomputer Center

University of California, San Diego

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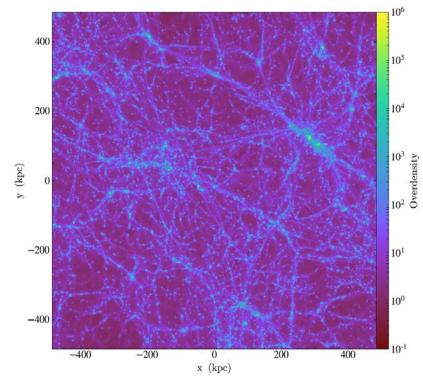


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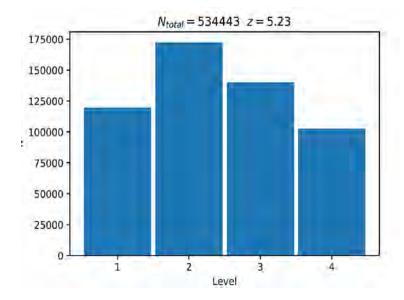
Science Motivation: Large-scale simulations of the IGM including galaxy feedback

- Precision comparisons between models and simulations require large volumes and high resolution in galaxies
- Standard ENZO code struggles to do this due to limited scalability of its AMR implementation (P<1000)
- In 2011, James Bordner and I embarked on a fromscratch redesign and reimplementation of ENZO capable of scaling AMR to millions of cores.
- Blue Waters has been instrumental in the development and testing of ENZO-E

PRAC: Realistic Simulations of the Intergalactic Medium: The Search for Missing Physics-Part 2



ENZO simulation



Projection of baryon density in a section of a 1024³, 25.6 Mpc box with 4 additional levels of refinement. 8K cores, Blue Waters #AMR grids versus level of refinement Potential level of concurrency: >100,000

Adopted Strategy

- Keep the best part of Enzo (numerical solvers) and replace the AMR infrastructure
- Implement using modern OOP best practices for modularity and extensibility
- Use the best available scalable AMR algorithm:
 - Array-of-Octrees (aka Forest-of-Octrees)
- Move from bulk synchronous to data-driven asynchronous execution model to support patch adaptive timestepping
- Leverage parallel runtimes that support this execution model, and have a path to exascale (Charm++)
- Make AMR software library application-independent so others can use it



Software Architecture

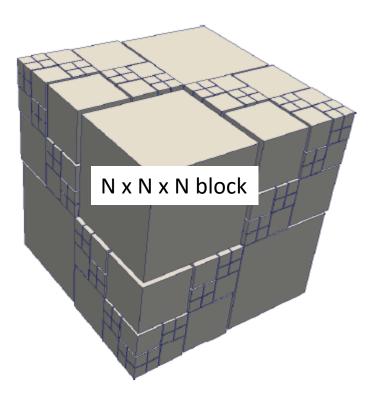
Enzo numerical solvers (Enzo-E)

Array-of-octrees AMR (Cello)

Charm++

Hardware (heterogeneous, hierarchical)

How does Cello implement AOT?



- Array of octrees of arbitrary size K x L x M
- An octree has leaf nodes which are **blocks** (N x N x N)
- Each block is a chare (unit of sequential work)
- The entire AOT is stored as a chare array using a bit indexing scheme
- Chare arrays are fully distributed data structures in Charm++

2 x 2 x 2 array

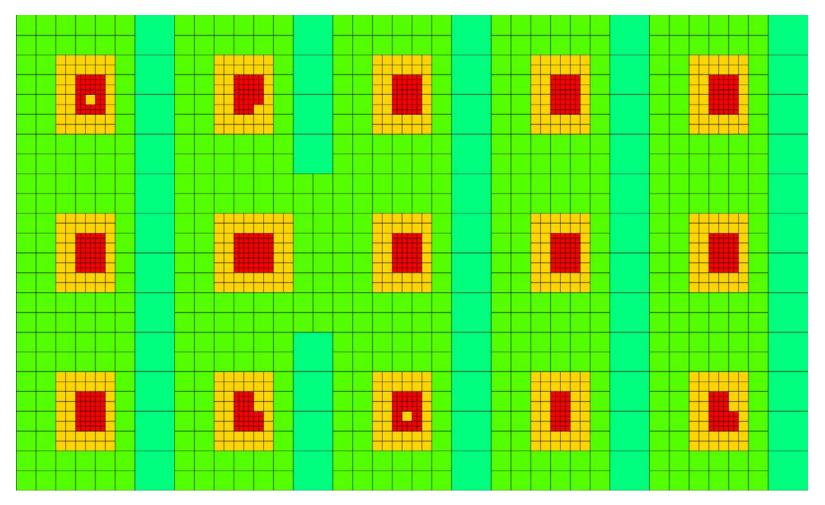
Demonstration of ENZO-E

Interacting blast waves with PPM solver - Total energy



Demonstration of ENZO-E

Mesh refinement level; 32³ blocks/chare



| Largest AMR |
|--------------------------------------|
| simulation |
| in the |
| world? |
| 1.7 trillion cells |
| 262K cores on NCSA Blue Waters |
| html |

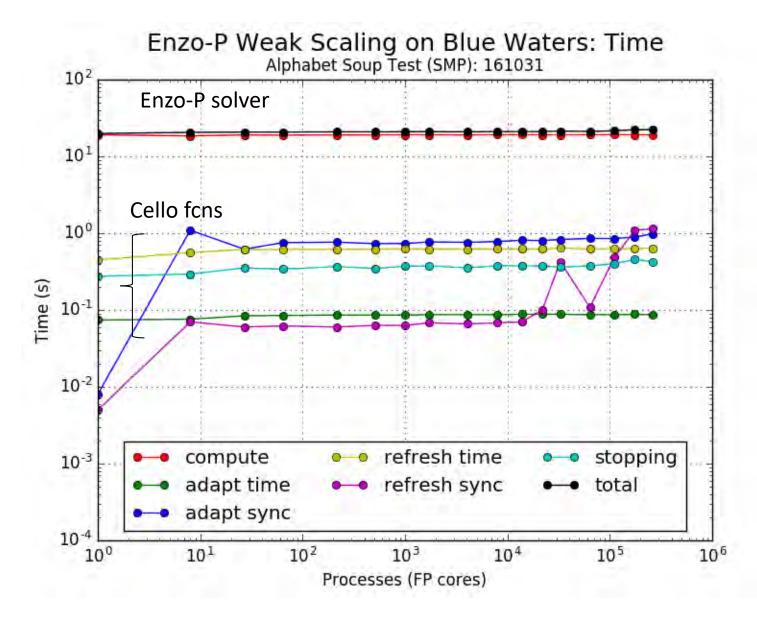
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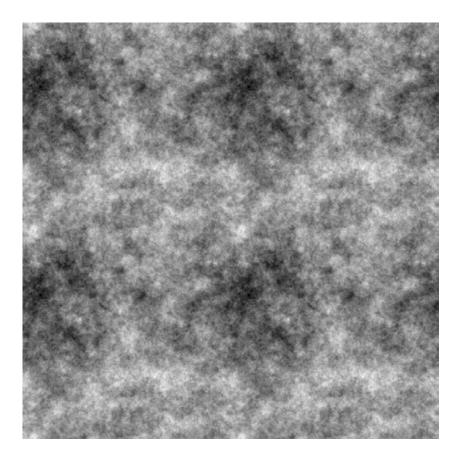
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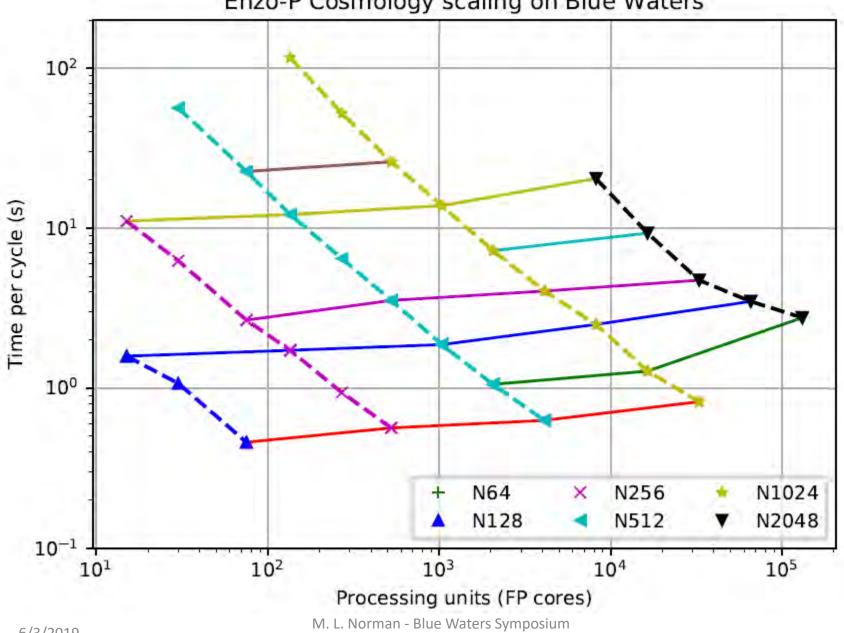


Hydrodynamic Cosmology Scaling Tests



Density projection in 512³ simulation

- Uniform grid only
- Weak and strong scaling
- 32³ blocks/chare
- 1, 8, 64 chares/core
- 64³ to 2048³ meshes
- p=8 to 128k cores

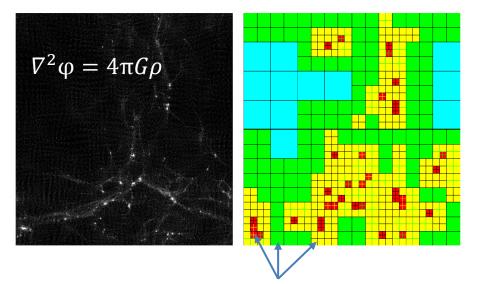


Enzo-P Cosmology scaling on Blue Waters

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Global Multilevel AMR Poisson Solver



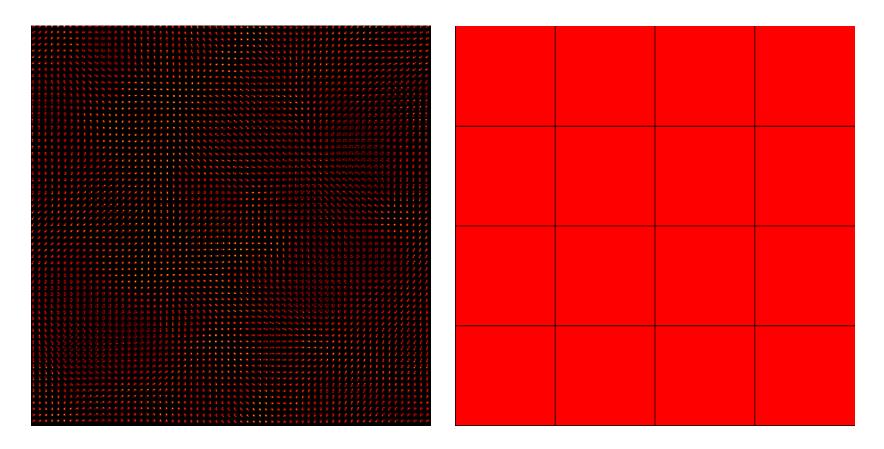
• Each square is projection of 16³ block

• Ax=b

- A is non-symmetric matrix arising from discretizing Laplacian operator on multilevel mesh
- x is gravitational potential $\boldsymbol{\varphi}$
- b is matter density source term
- Algorithm
 - BiCGStab
 - Diagonally-preconditioned
 - Multigrid-preconditioned
 - Parallelize over all blocks in AOT using Charm++
- poor scalability

AMR cosmology with HG* solver

64³ mesh (4³ array of 16³ blocks), 4 AMR levels (1024³ eff.), PE=8



* HG = multigrid preconditioned BiCGStab

AMR Cosmology Blocks per Level N=64³ P=8 1000 ~1000 blocks/PE 100 at all levels 10 1 0.1 level 0 level 1 0.01 level 2 level 3 level 4 0.001 100 50 150 200 250 300 350 400 0 Cycle

Blocks / PE / Level

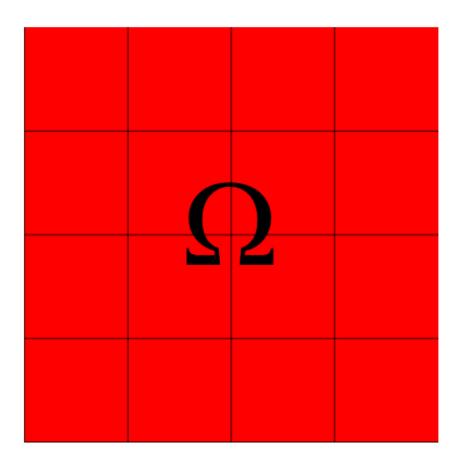
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Domain-decomposed AMR Poisson Solver (DD)

| | Ω | Ω | Ω | Ω |
|---|---|---|---|---|
| | Ω | Ω | Ω | Ω |
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Each block becomes an octree

Domain-decomposed AMR Poisson Solver (DD)



- Step 1: project density field to root grid blocks
- Step 2: global Poisson solve on root grid using multigrid solver
- Step 3: interpolate Φ_0 to faces of each octree

Each block becomes an octree

Domain-decomposed AMR Poisson Solver (DD)

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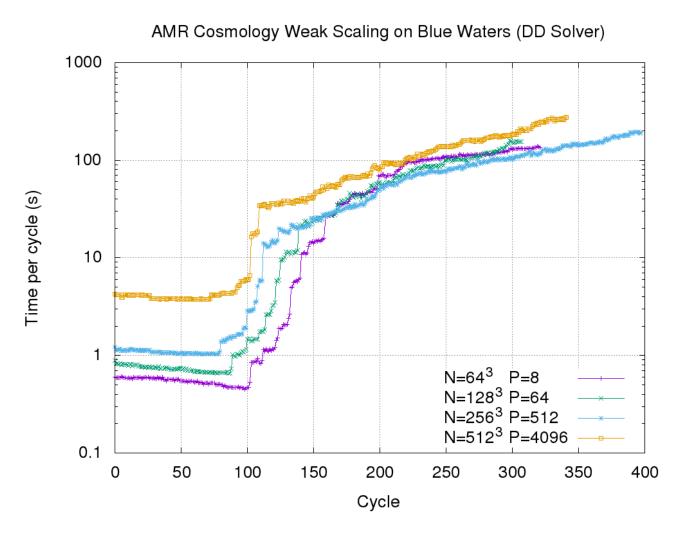
- Step 4: local Poisson solve on each octree using BiCGtab
- Step 5: Jacobi smooth potential Φ_{ℓ} on leaf blocks

Each block becomes an octree

DD in action

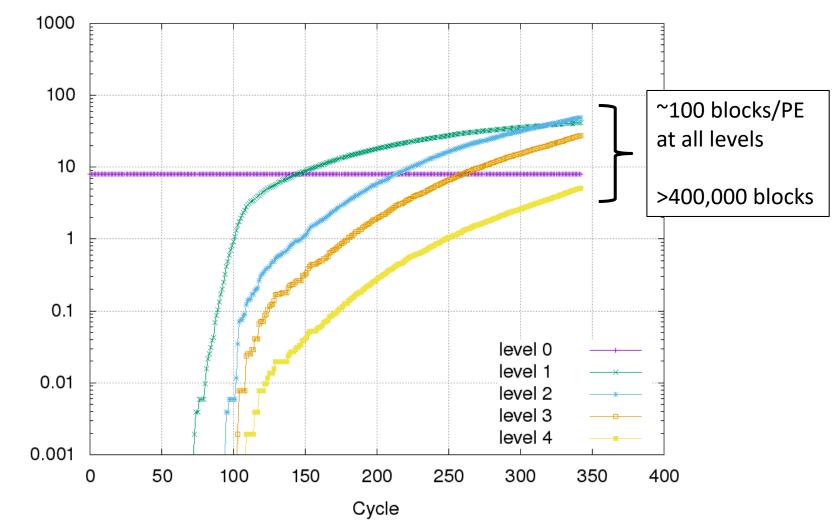
128³ mesh (8³ array of 16³ blocks), 4 AMR levels (2048³ eff.), PE=64

Is DD scalable? Yes!





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AMR Cosmology Blocks per Level N=512³ P=4096

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NSF CSSI grant (ENZO-E)

- \$1.9M, 3 years
- Goal: feature-complete implemention of ENZO solvers into ENZO-E
- Goal: migration of ENZO community to ENZO-E
- Goal: implementations for exascale (accelerators)
- PI team
 - Mike Norman, James Bordner (UCSD)
 - Brian O'Shea (MSU)
 - Greg Bryan (Columbia)
 - John Wise (Georgia Tech)

ENZO-E tasks

- FMM gravity solver
- Block-adaptive local timestepping
- Adaptive ray tracing radiative transfer
- Cosmic ray transport incl. MHD
- Interfacing to GRACKLE chemistry library
- Interface to GPUs using Kokkos
- Lots of scaling/optimization work......

Contribution of Blue Waters

- You need a petascale platform to develop an exascale code!
 - Sheer size
 - Balanced architecture
 - Throughput (especially scaling runs)
 - Favorable Q policies
 - Mature SW environment
- Thanks for the memories!!!!