

### A Scalable AMR Gravity Solver for ENZO-E (Extreme-scale ENZO)

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Science Motivation: Large-scale simulations of the IGM including galaxy feedback

- Precision comparisons between models and simulations require large volumes and high resolution in galaxies
- Standard ENZO code struggles to do this due to limited scalability of its AMR implementation (P<1000)
- In 2011, James Bordner and I embarked on a from- scratch redesign and reimplementation of ENZO capable of scaling AMR to millions of cores.
- Blue Waters has been instrumental in the development and testing of ENZO-E

### PRAC: Realistic Simulations of the Intergalactic Medium: The Search for Missing Physics-Part 2



ENZO simulation



Projection of baryon density in a section of a  $1024^3$ , 25.6 Mpc box with 4 additional levels of refinement. 8K cores, Blue Waters

#AMR grids versus level of refinement Potential level of concurrency: >100,000

# Adopted Strategy

- Keep the best part of Enzo (numerical solvers) and replace the AMR infrastructure
- Implement using modern OOP best practices for modularity and extensibility
- Use the best available scalable AMR algorithm:
	- Array-of-Octrees (aka Forest-of-Octrees)
- Move from bulk synchronous to data-driven asynchronous execution model to support patch adaptive timestepping
- Leverage parallel runtimes that support this execution model, and have a path to exascale (Charm++)
- Make AMR software library application-independent so others can use it



## **Software Architecture**

Enzo numerical solvers (Enzo-E)

Array-of-octrees AMR (Cello)

Charm++

Hardware (heterogeneous, hierarchical)

# How does Cello implement AOT?



- **Array of octrees** of arbitrary size  $K \times I \times M$
- An octree has leaf nodes which are **blocks** (N x N x N)
- Each block is a **chare** (unit of sequential work)
- The entire AOT is stored as a **chare array** using a bit indexing scheme
- Chare arrays are **fully distributed data structures** in Charm++

2 x 2 x 2 array

## Demonstration of ENZO-E

Interacting blast waves with PPM solver - Total energy



## Demonstration of ENZO-E

Mesh refinement level; 32<sup>3</sup> blocks/chare





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## Hydrodynamic Cosmology Scaling Tests



### Density projection in 512<sup>3</sup> simulation

- Uniform grid only
- Weak and strong scaling
- 32<sup>3</sup> blocks/chare
- 1, 8, 64 chares/core
- $64^3$  to 2048<sup>3</sup> meshes
- p=8 to 128k cores



### Enzo-P Cosmology scaling on Blue Waters

## Global Multilevel AMR Poisson Solver



Each square is projection of 16<sup>3</sup> block

 $Ax = b$ 

- A is non-symmetric matrix arising from discretizing Laplacian operator on multilevel mesh
- $-$  x is gravitational potential  $\phi$
- b is matter density source term
- Algorithm
	- BiCGStab
		- Diagonally-preconditioned
		- Multigrid-preconditioned
	- Parallelize over all blocks in AOT using Charm++
- poor scalability

# AMR cosmology with HG\* solver

64<sup>3</sup> mesh (4<sup>3</sup> array of 16<sup>3</sup> blocks), 4 AMR levels (1024<sup>3</sup> eff.), PE=8



#### \* HG = multigrid preconditioned BiCGStab



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### Domain-decomposed AMR Poisson Solver (DD)



#### Each block becomes an octree

### Domain-decomposed AMR Poisson Solver (DD)



- Step 1: project density field to root grid blocks
- Step 2: global Poisson solve on root grid using multigrid solver
- Step 3: interpolate  $\Phi_0$  to faces of each octree

#### Each block becomes an octree

### Domain-decomposed AMR Poisson Solver (DD)



- Step 4: local Poisson solve on each octree using BiCGtab
- Step 5: Jacobi smooth potential  $Φ$ <sub>*l*</sub> on leaf blocks

#### Each block becomes an octree

## DD in action

128<sup>3</sup> mesh (8<sup>3</sup> array of 16<sup>3</sup> blocks), 4 AMR levels (2048<sup>3</sup> eff.), PE=64



## Is DD scalable? Yes!





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#### AMR Cosmology Blocks per Level N=512<sup>3</sup> P=4096

Blocks / PE / Level

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# **NSF CSSI grant (ENZO-E)**

- $\bullet$  \$1.9M, 3 years
- Goal: feature-complete implemention of ENZO solvers into ENZO-E
- Goal: migration of ENZO community to ENZO-E
- Goal: implementations for exascale (accelerators)
- PI team
	- Mike Norman, James Bordner (UCSD)
	- Brian O'Shea (MSU)
	- Greg Bryan (Columbia)
	- John Wise (Georgia Tech)

## ENZO-E tasks

- FMM gravity solver
- Block-adaptive local timestepping
- Adaptive ray tracing radiative transfer
- Cosmic ray transport incl. MHD
- Interfacing to GRACKLE chemistry library
- Interface to GPUs using Kokkos
- Lots of scaling/optimization work.......

# Contribution of Blue Waters

- You need a petascale platform to develop an exascale code!
	- Sheer size
	- Balanced architecture
	- Throughput (especially scaling runs)
	- Favorable Q policies
	- Mature SW environment
- Thanks for the memories!!!!