

Simulations of the Structure of Magnetic Fields in Galaxy Clusters

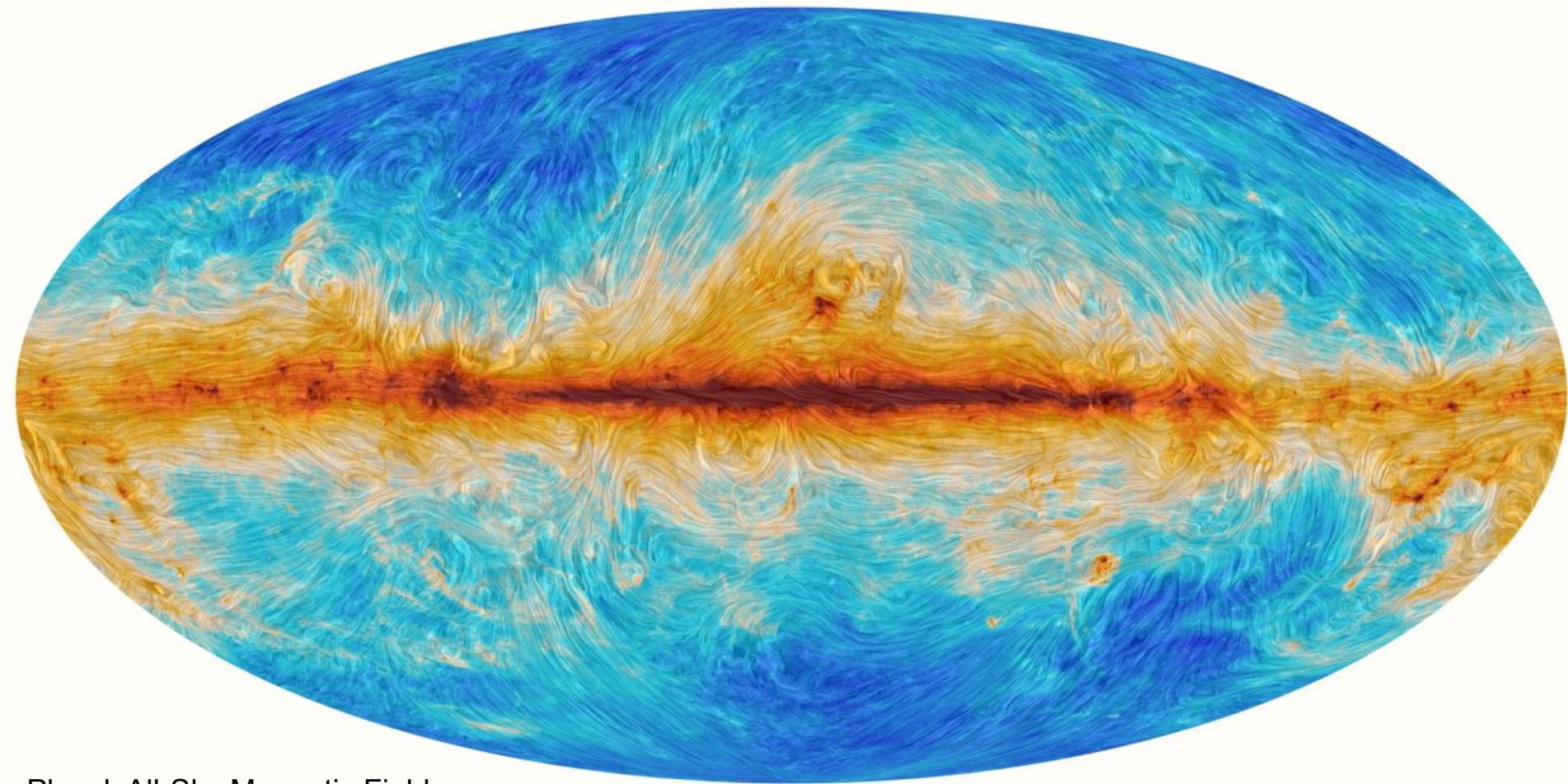
Forrest W. Glines^{1,2},
Brian W. O'Shea^{1,2}, Philipp Grete¹



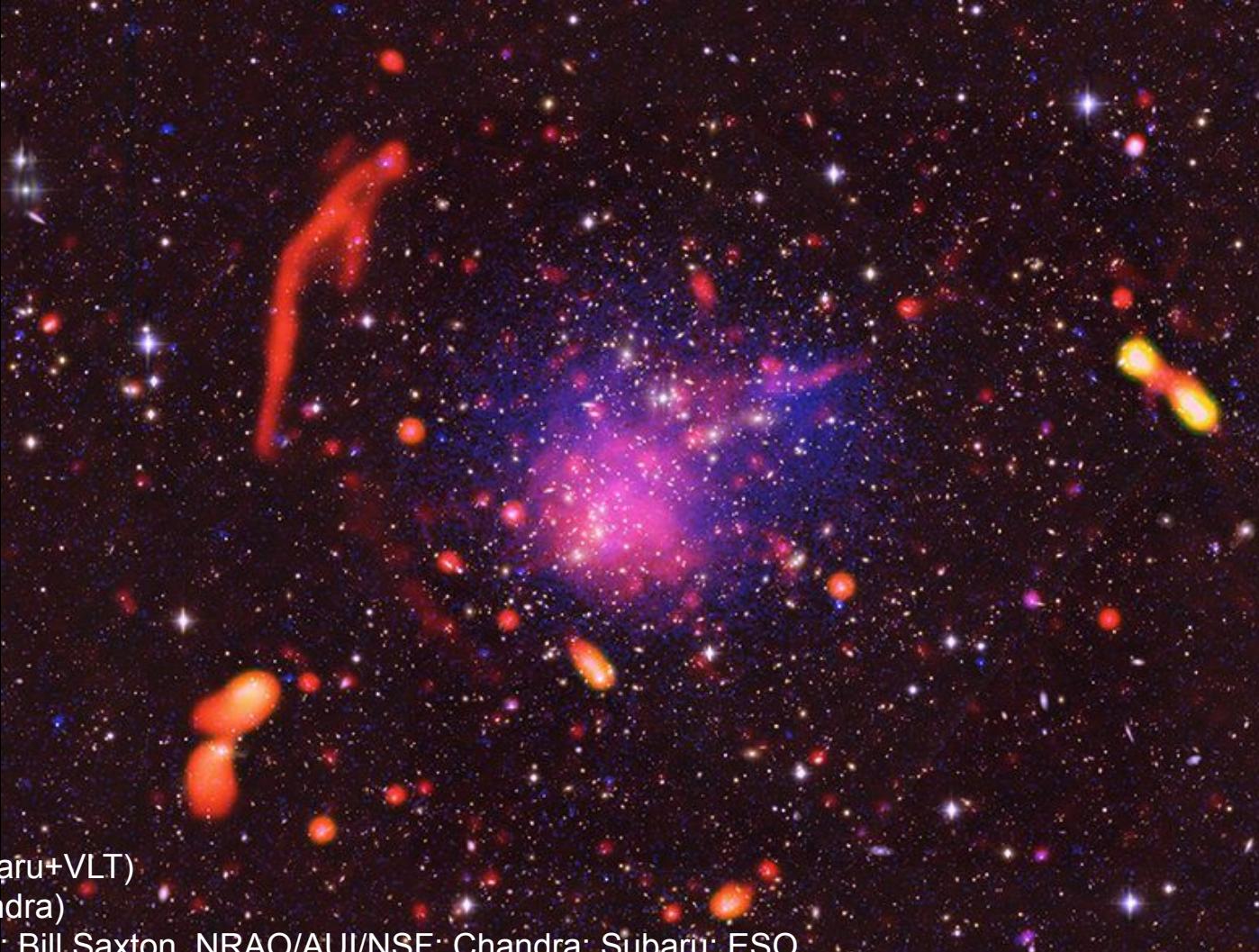
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Planck All-Sky Magnetic Field



Abell 2744

Radio (VLA)

Optical (Subaru+VLT)

X-Ray (Chandra)

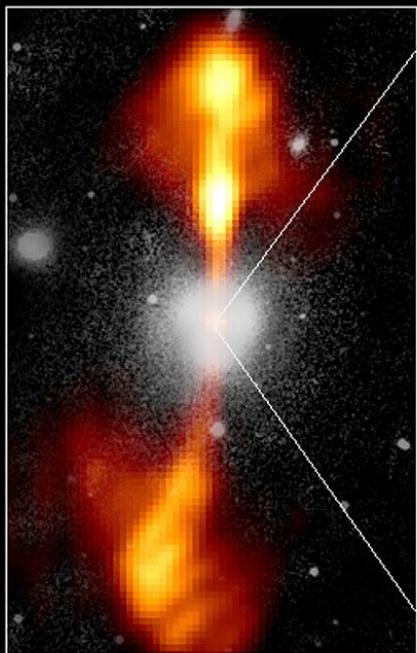
Pearce et al.; Bill Saxton, NRAO/AUI/NSF; Chandra; Subaru; ESO.

Core of Galaxy NGC 4261

Hubble Space Telescope

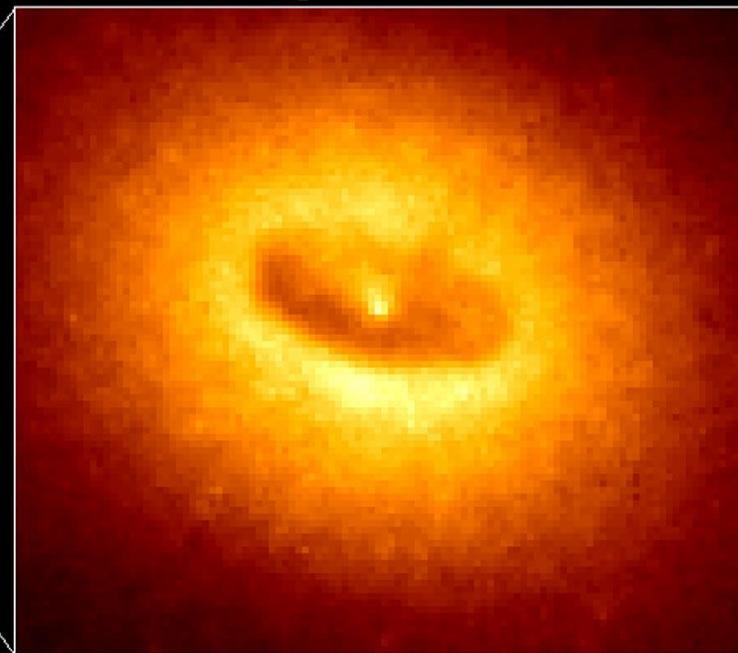
Wide Field / Planetary Camera

Ground-Based Optical/Radio Image



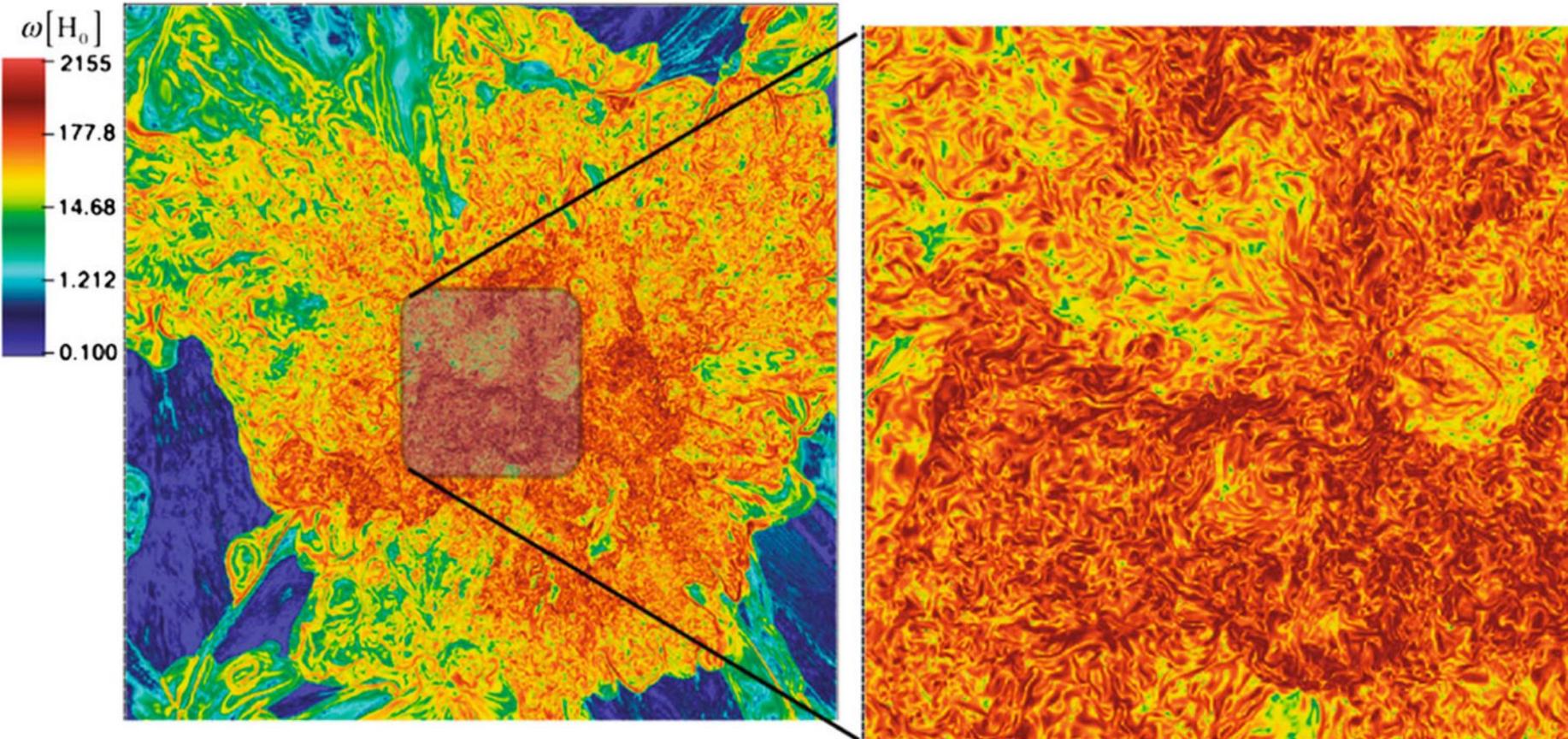
380 Arc Seconds
88,000 LIGHTYEARS

HST Image of a Gas and Dust Disk



1.7 Arc Seconds
400 LIGHTYEARS

NASA



Minati 2014

CPU



Xeon Phi



GPU



AMD



NVIDIA.

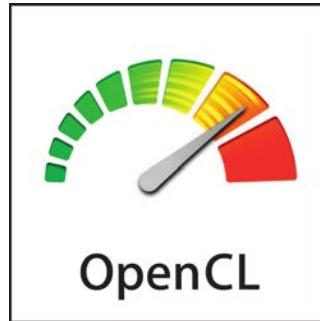
BLUE WATERS
SUSTAINED PETASCALE COMPUTING



Performance Portability APIs

Typical Solutions

- OpenCL
- OpenACC
- OpenMP 4.5, 5.0



Abstraction Layers

- OCCA
- RAJA
- Kokkos

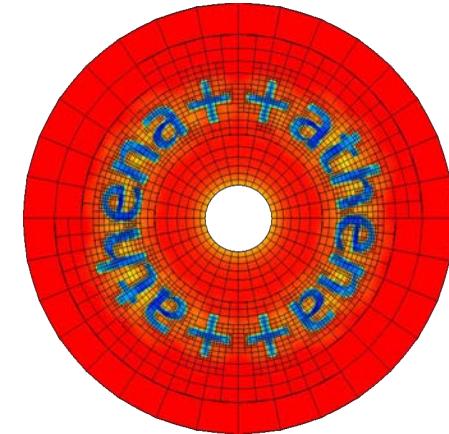


K-Athena: Performance Portable MHD

- Partial Conversion of astrophysical MHD code Athena++ using Kokkos library
- Minimal code changes
- Performance Portable
 - Ready for CPUs, GPUs
- Freely available on GitLab

<https://gitlab.com/pgrete/kathena>

<https://arxiv.org/abs/1905.04341>



Kokkos Conversion

Athena++

```
for( int k = ks; k < ke; k++) {  
    for( int j = js; j < je; j++) {  
        #pragma omp simd  
        for( int i = is; i < ie; i++) {  
            /* Loop Body */  
            u(k,j,i) = ...  
        } } }
```

K-Athena

```
using namespace Kokkos;  
parallel_for( MDRangePolicy<Rank<3>>  
    ({ks,js,is}, {ke,je,ie}),  
    KOKKOS_LAMBDA(int k, int j, int i){  
        /* Loop Body */  
        u(k,j,i) = ...  
    });
```

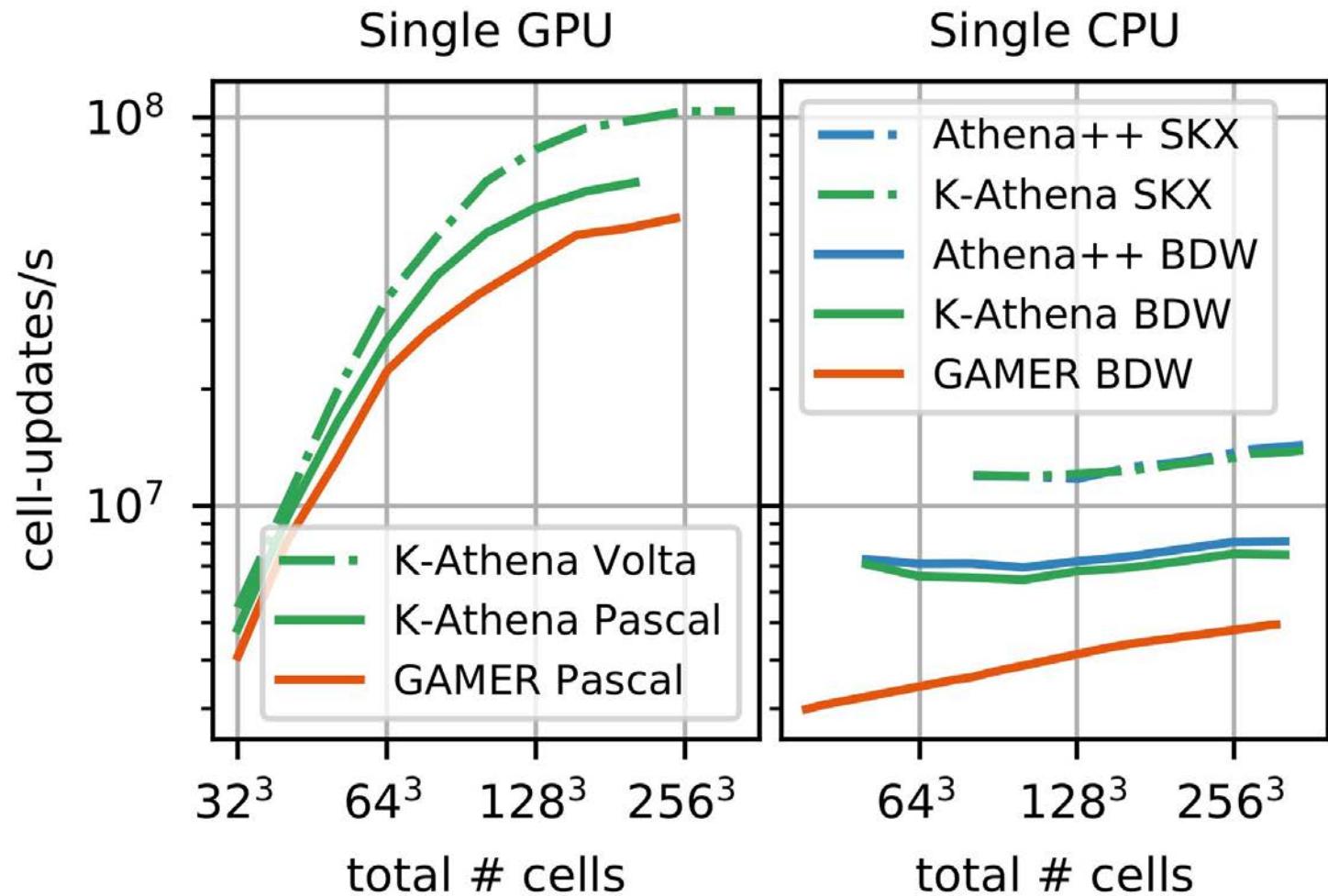
Parallelism Flexibility

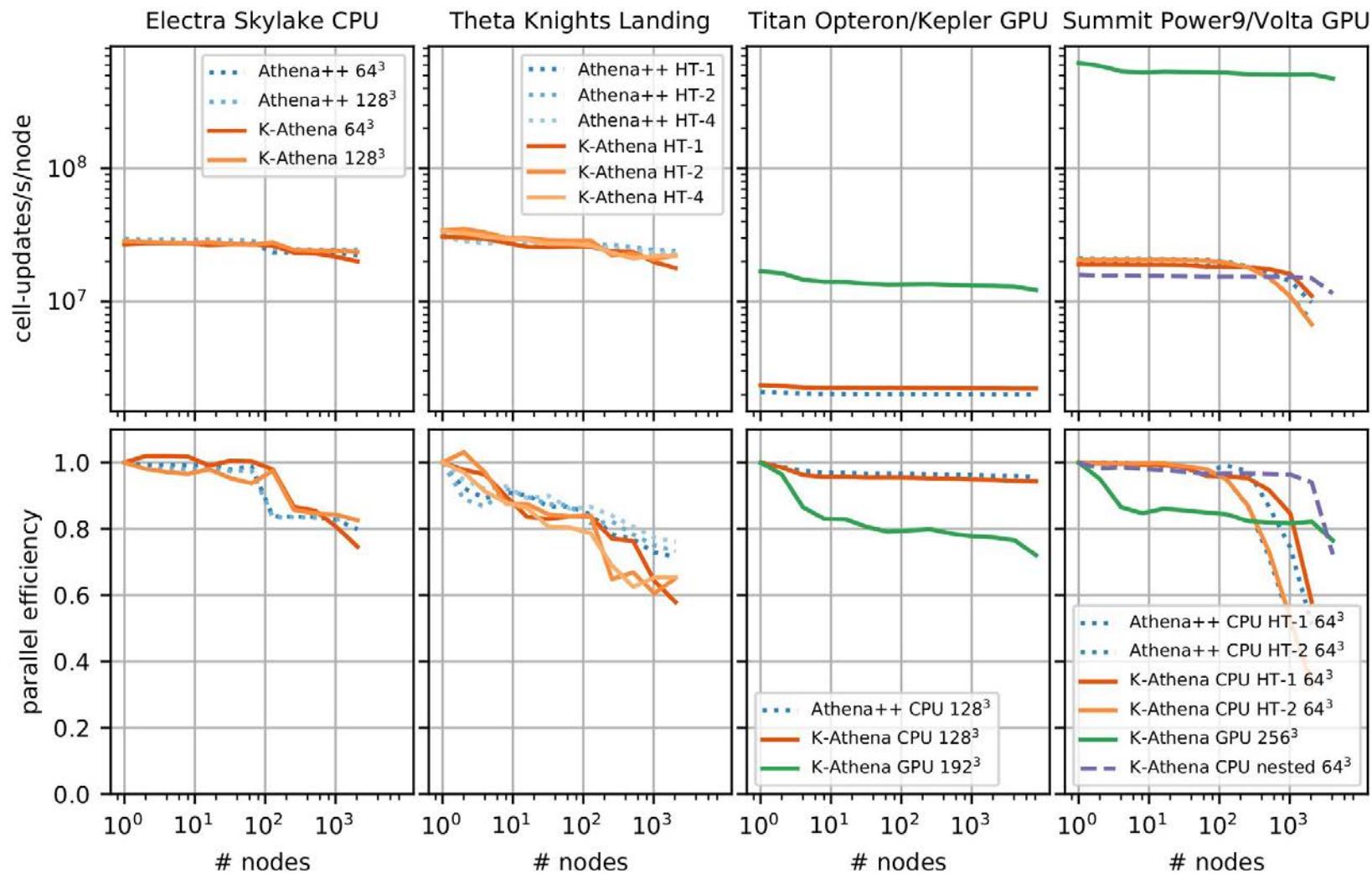
Flat Parallelism

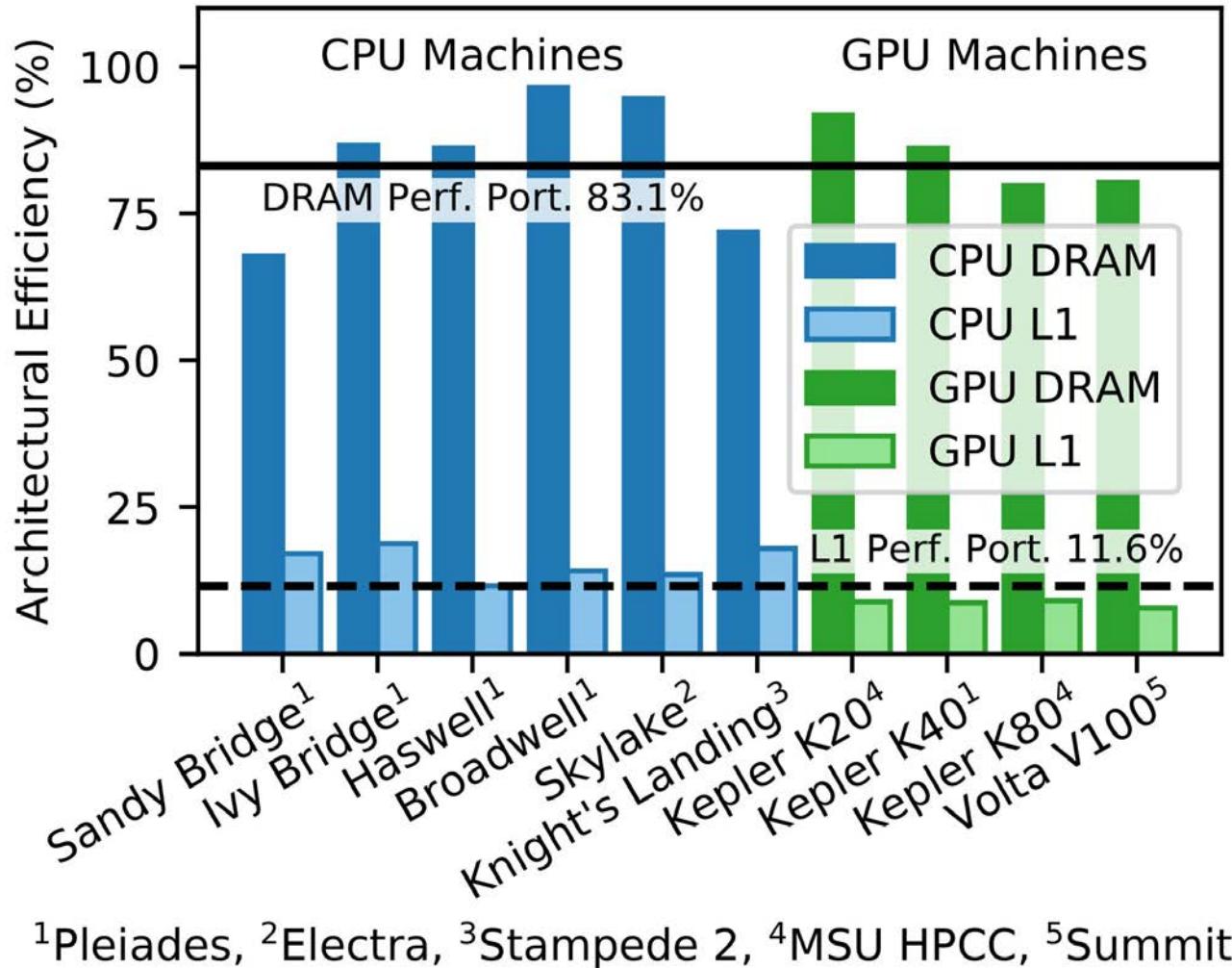
```
using namespace Kokkos;
parallel_for( MDRangePolicy<Rank<3>>
    ({ks,js,is},{ke,je,ie}),
    KOKKOS_LAMBDA(int k, int j, int i){
        /* Loop Body */
        u(k,j,i) = ...
});
```

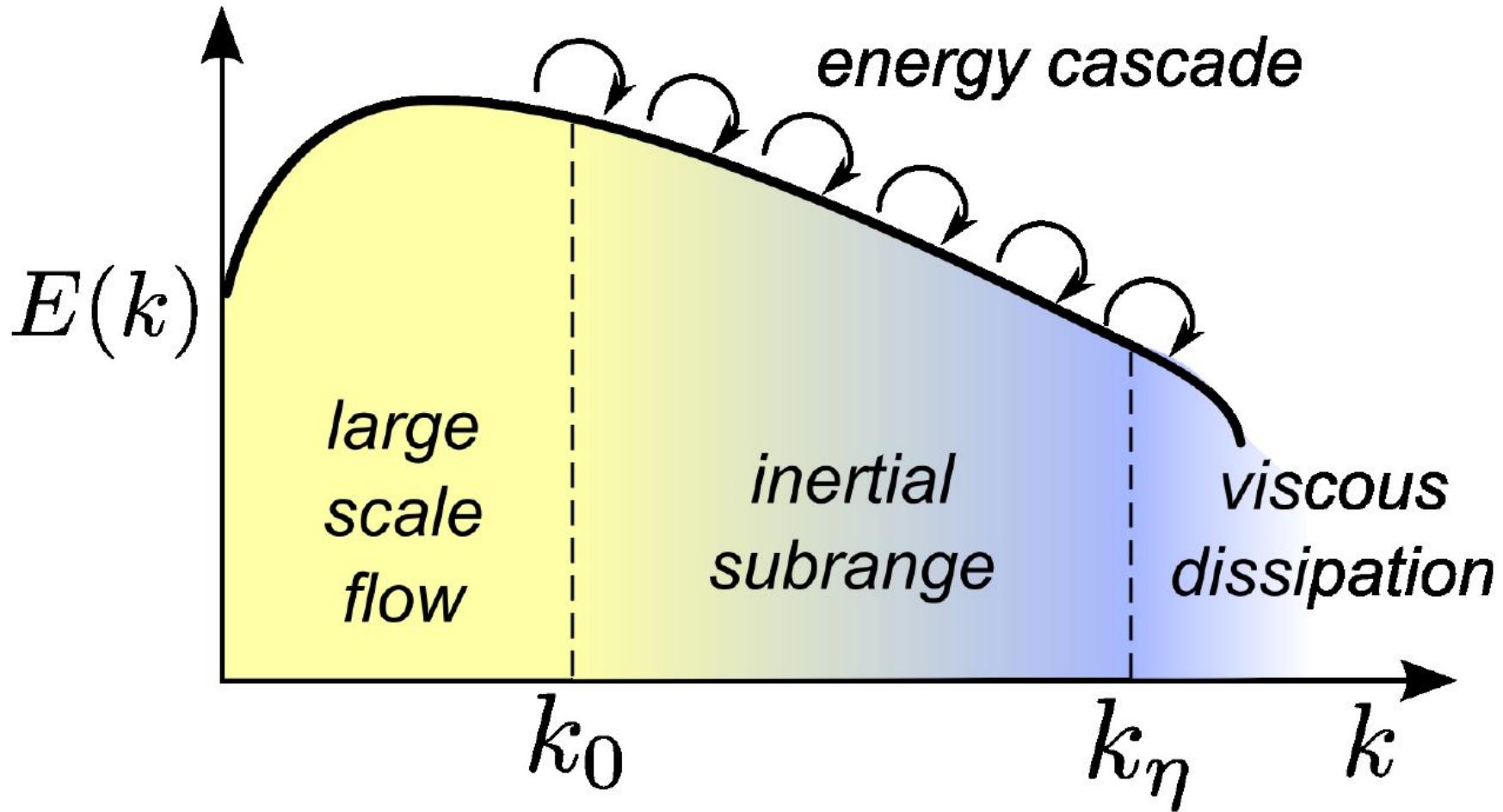
Hierarchical Parallelism

```
using namespace Kokkos;
parallel_for(team_policy(nk*nj, AUTO),
    KOKKOS_LAMBDA(member_type team_mem) {
        int lr = team_mem.league_rank();
        int k = lr / nj + ks;
        int j = lr % nj + js;
        parallel_for(
            TeamThreadRange<>(team_mem,is,ie),
            [&](int i) {
                /* Loop Body */
                u(k,j,i) = ...
            });
});
```

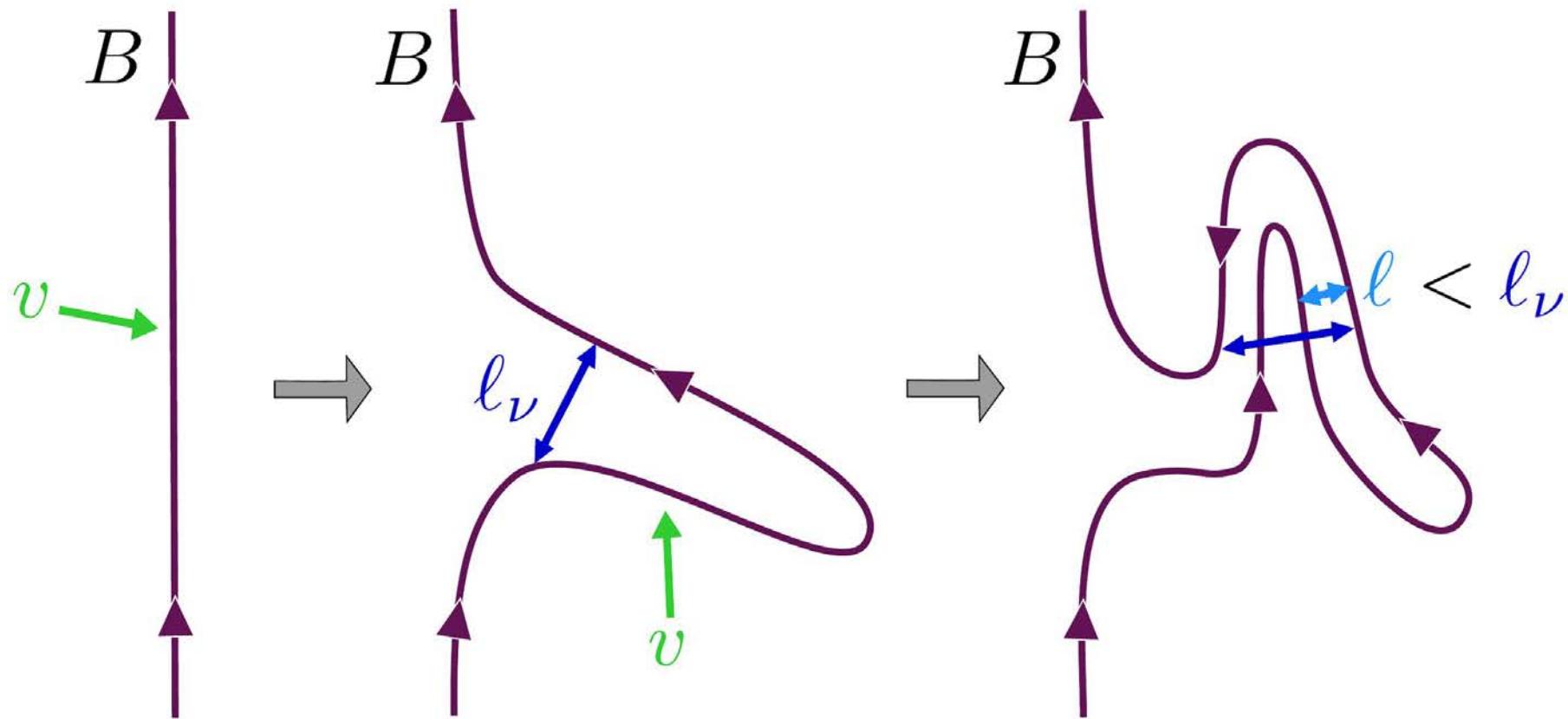


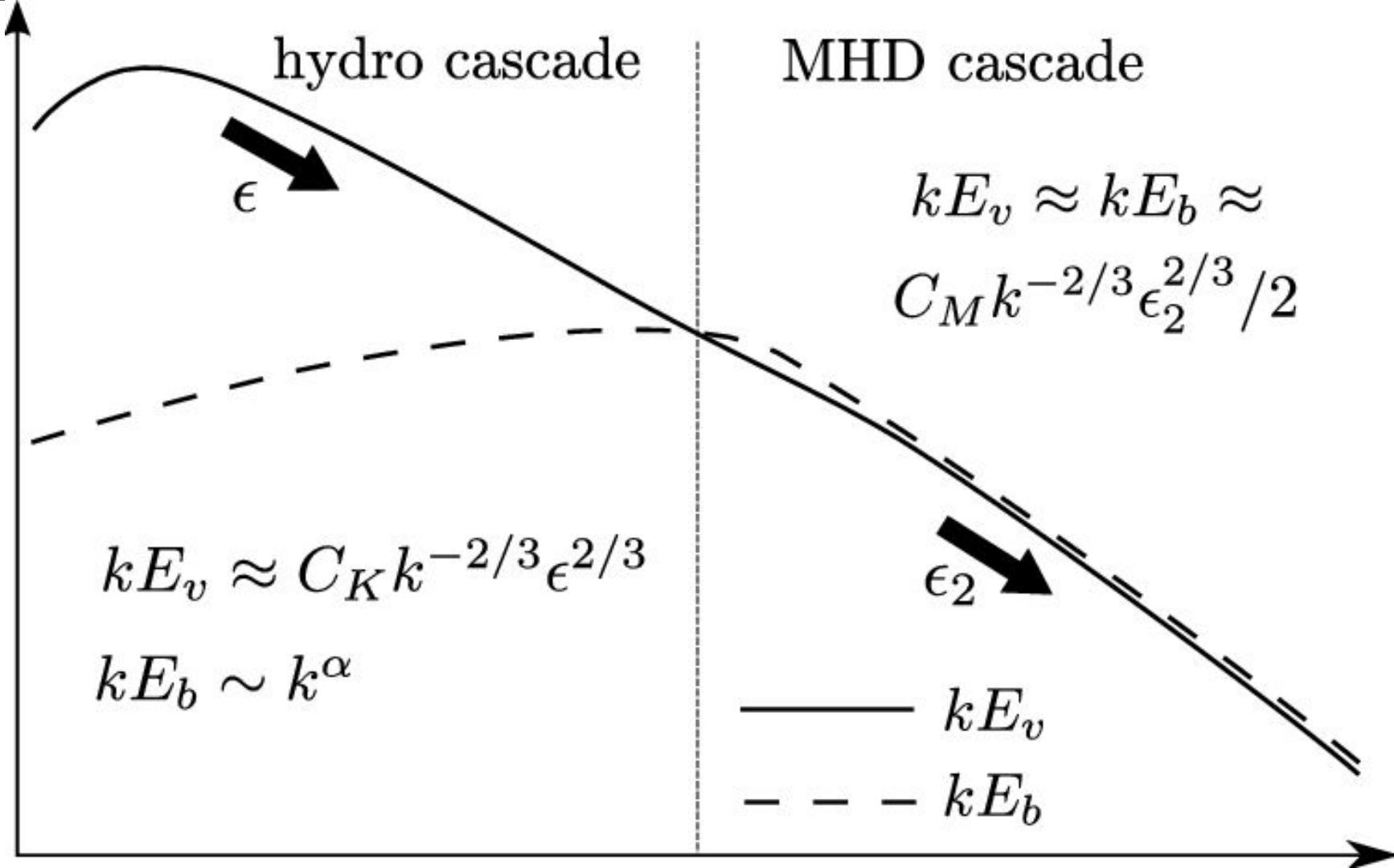


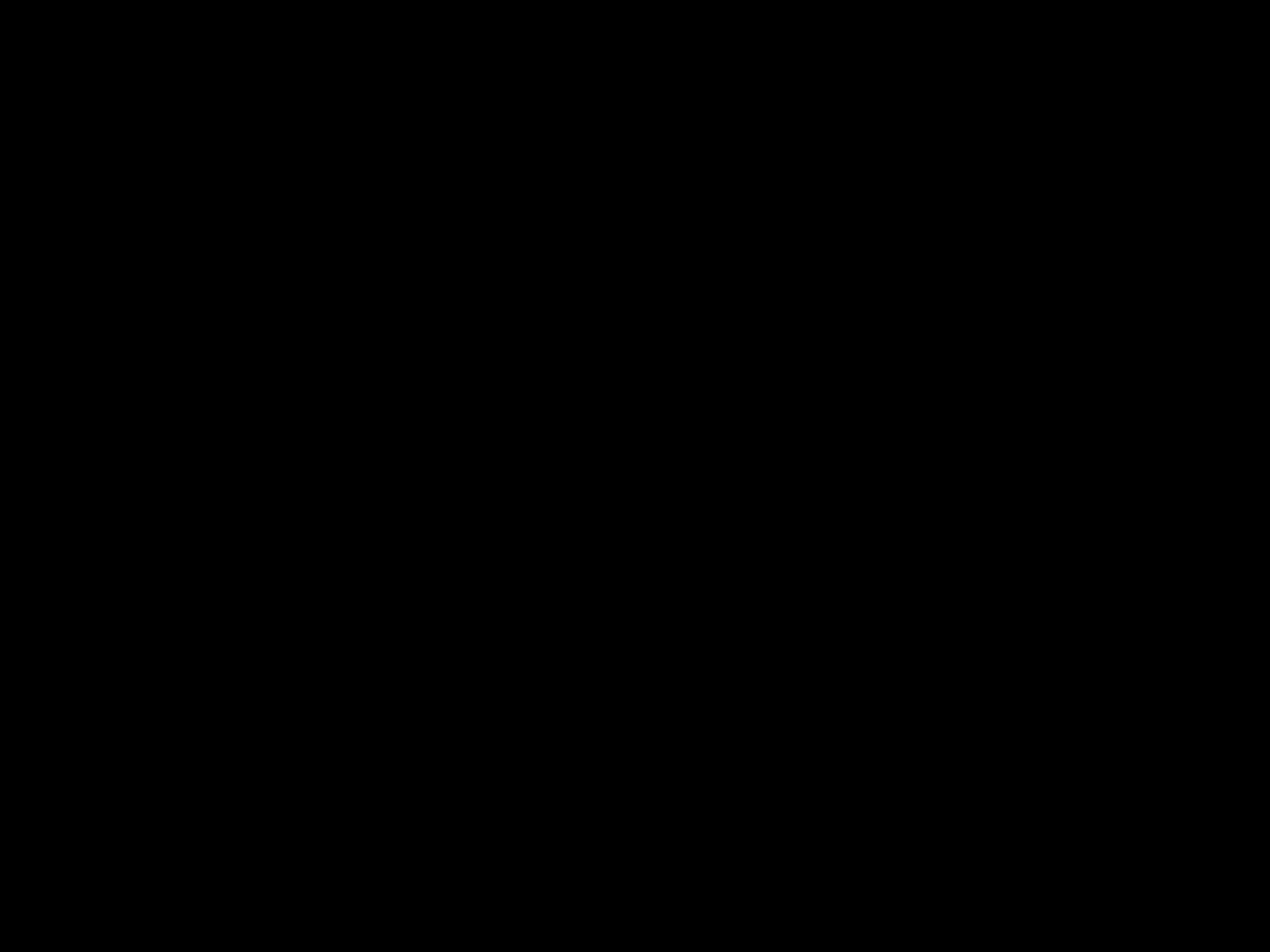




Small Scale Turbulent Dynamo

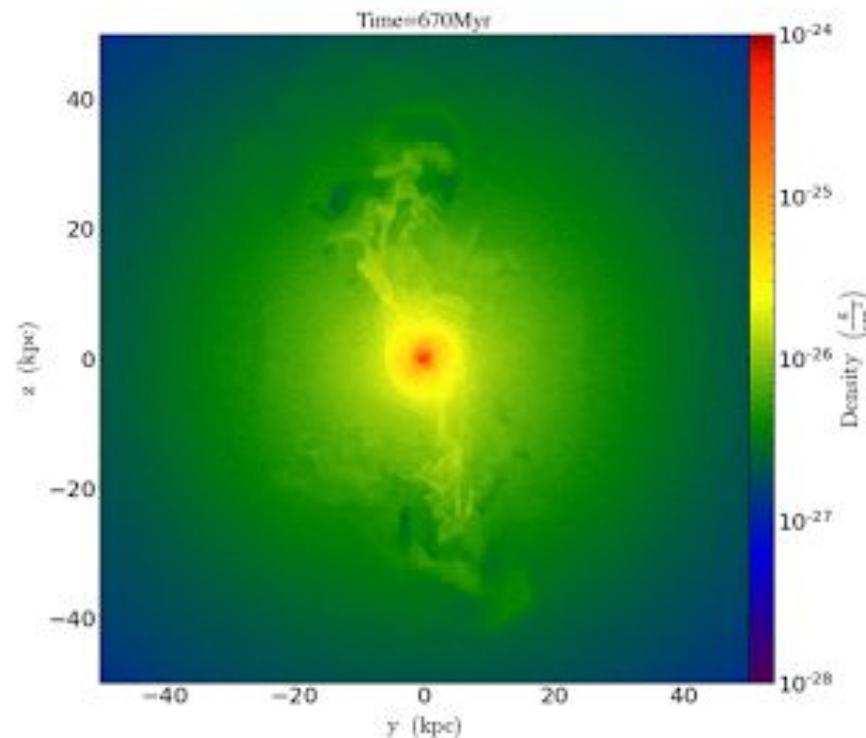






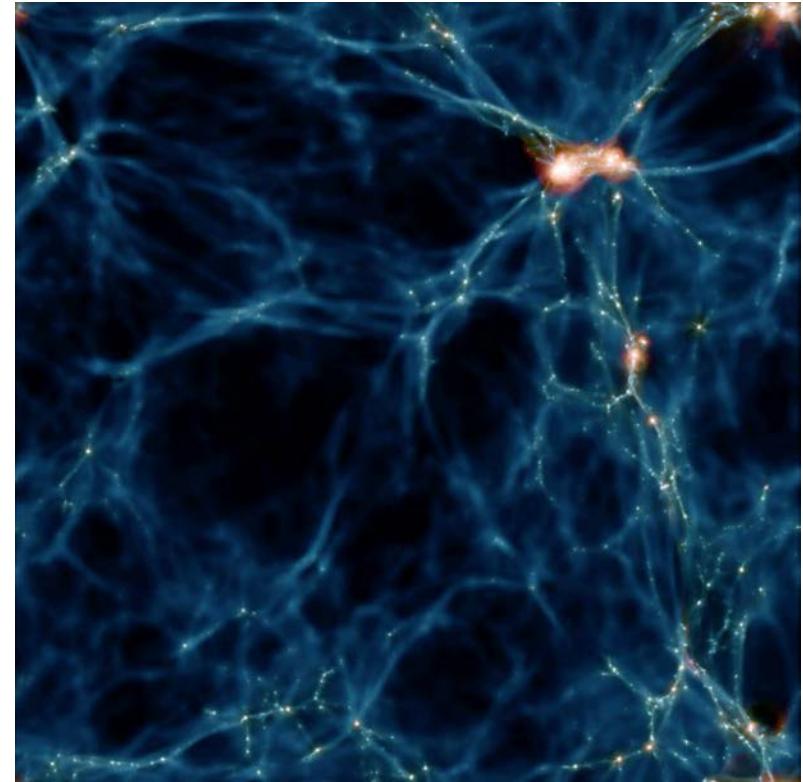
Isolated Galaxy Cluster Simulations with K-Athena

- Physics rich cluster simulation
 - MHD
 - Magnetic AGN feedback
 - Tabulated Cooling
 - Cosmic Ray model
 - Viscosity and Conduction
- How does the ISM behave?
- How does AGN feedback thermalize?



Cosmological Simulations with Enzo-E

- How does rich plasma physics change cosmological simulations?
- How do cool core cluster emerge, and how are they maintained?
- How do AGN affect star formation and the evolution of galaxies within the cluster?
- With all the plasma physics, do we get realistic cool core clusters?



Romulus C Simulation [Tremmel 2018]

References

K-Athena: <https://gitlab.com/pgrete/kathena>

References

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