



# Three-Dimensional Models of Core-Collapse Supernova Explosions

Adam Burrows, Davíd Vartanyan, Davíd Radíce, Hírokí Nagakura, Víktoríya Morozova, Aaron Skínner, Josh Dolence Supported by: DOE/SciDAC4 NSF/MPPC NSF/AST BlueWaters; INCITE; XSEDE

### Essential Elements of Neutrino Mechanism

- Pseudo-Chandrasekhar core collapses for hundreds of seconds
- Bounces at nuclear densities and launches a shock wave
- Shock wave stalls due to breakout neutrino losses and photodissociation of accreta within 10's of milliseconds at ~100-150 km into an accretion shock
- Neutrino emission from the inner core (PNS) heats the "gain region" behind the shock, and drives turbulent convection
- Neutrino energy deposition behind the shock and turbulent pressure together eventually overcome the ram pressure of the continuing accretion to launch a supernova
- Delayed Explosion
- Core-collapse supernova explosion is a critical phenomenon/ bifurcation between steady solutions and exploding solutions
- Multi-D (expensive) necessary because most models don't explode (aren't reenergized) in 1D (spherical), but require the extra turbulent pressure/stress of neutrino-driven convection (and other effects)

### Core-Collapse Theory: What's New?

- Turbulence crucial to most explosions, necessitating multi-D treatment
- In the last ten years, we could do multiple 2D simulations every year to explore parameters, understand systematics, and explore progenitor structure dependence.
- Techniques improved and computers sped up; resolution-dependence
- Can now do multiple 3D simulations per year (and afford to make a few mistakes!)
- GR, Many-body neutrino-matter corrections (more to do), and PNS convection lead to enhanced  $v_{\mu}$  losses, faster contraction, hence hotter  $v_e$  and anti- $v_e$  neutrinospheres
- Incorporated inelastic neutrino-matter processes extra neutrino heating
- Accretion of the Si/O interface; seed perturbations of progenitor (?)

### FORNAX: 1D,2D,3D, Multi-Group, Radiation/Hydrodynamics

### FORNAX: 1D,2D,3D, Multi-Group, Explicit Radiation/Hydrodynamics

- Solves the Two-Moment Transport Equations, with 2<sup>nd</sup> and 3<sup>rd</sup> moment closures (not "ray-by-ray"); second-order accurate in space and time
- Explicit Riemann Godunov-like solution to the Transport operator
- Terms of O(v/c) included in transport; inelastic/redistribution scattering
- Implicit solution to the local transport source terms
- Explicit hydro; full energy and momentum couplings HLLC
- Conserves energy and momentum to machine precision
- Very good energy conservation with gravity included
- "6"– Dim. = 1(time) + 3(space) + 1(energy-group) + vector Flux
- Logically spherical coordinates general metric/covariant formulation
- Multipole Gravity (includes GR-like modifications to the monopole)
- Multi-D calculated to the center Core refinement ("dendritic grid") improves timestepping by many factors (!); static mesh refinement
- Good strong scaling in core count and scaling in energy group
- Result: Fast multi-D supernova code (by factor of ~5-10 x many other codes)
- Skinner et al. 2016 ; Radice et al. 2017; Burrows et al. 2018; Skinner et al. 2019; Burrows et al. 2019; Vartanyan et al. 2018,2019; Nagakura et al.

### FORNAX (cont.)

- Includes: Inelastic scattering off electrons
- Inelastic scattering off nucleons
- Includes in-medium Many-body response corrections (Horowitz et al. 2017)
- General-relativistic monopole gravity correction and gravitational redshifts (can compare with Newtonian)
- Multi-D transport, with rbr+ option (for comparison)
- Weak magnetism and recoil corrections

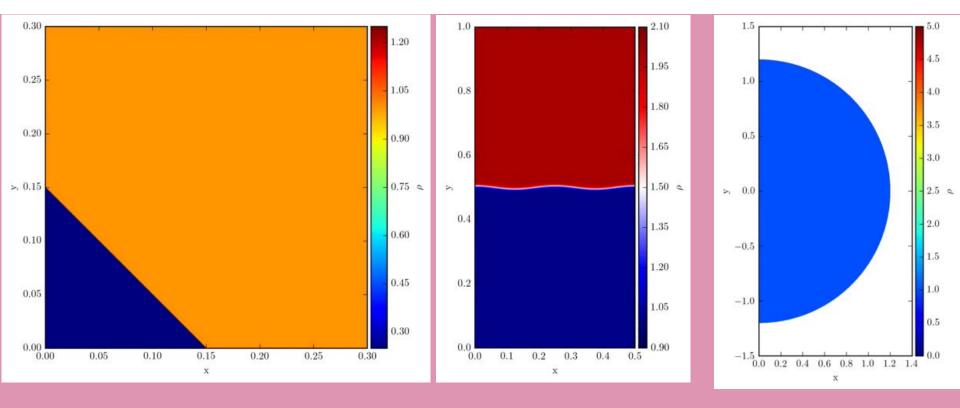
#### **Fornax Papers**

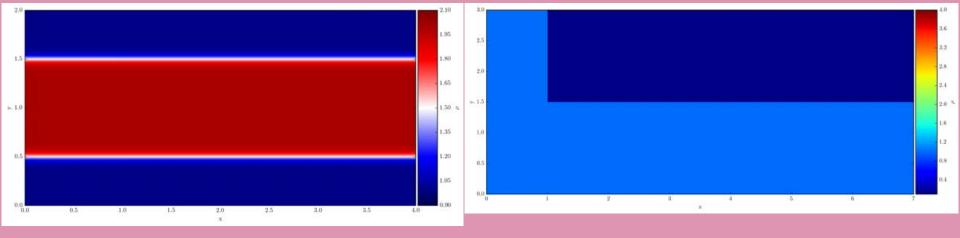
- Wallace et al. 2016 Neutrino breakout signal
- Skinner et al. 2016 Ray-by-ray+ study
- Radice et al. 2017 Electron-capture supernovae
- Burrows et al. 2018 Crucial component study
- Morozova et al. 2018 Gravitational wave signal (2D)
- Vartanyan et al. 2018 "Revival of the fittest"
- Seadrow et al. 2018 Signals in neutrino detectors
- O'Connor et al. 2018 1D code comparison
- Skinner et al. 2019 Fornax code paper
- Radice et al. 2019 Gravitational waves (3D)
- Vartanyan et al. 2019 3D explosion model
- Burrows et al. 2019 Multiple low-mass 3D explosion models
- Nagakura et al. 2019 3D model Resolution study

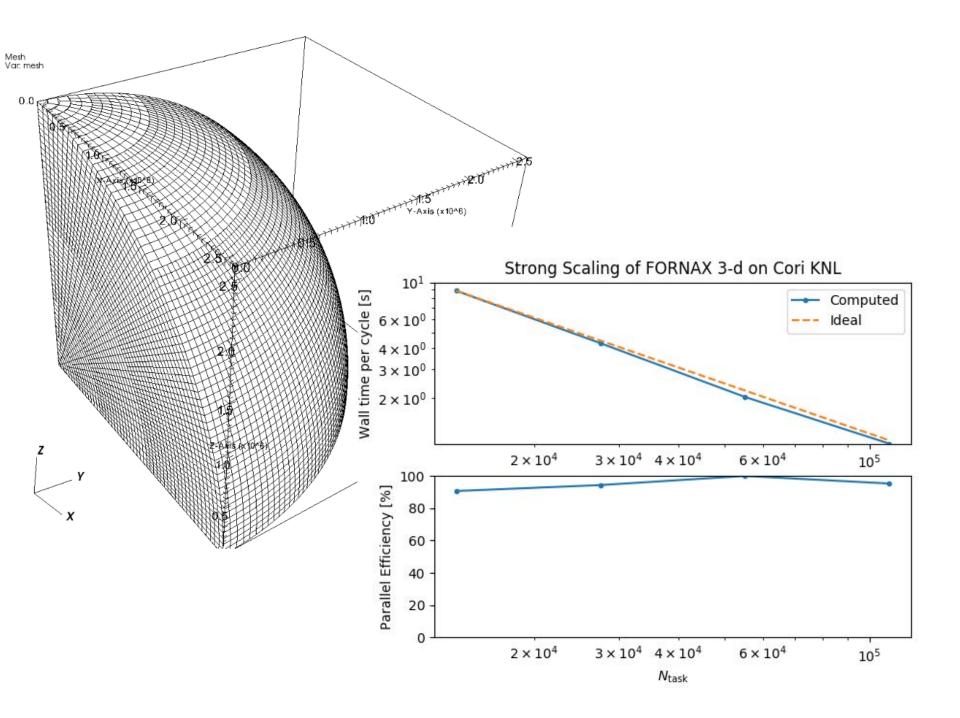
Recent 3D Fornax Simulations with Necessary Realism

- 9, 10, 11, 12, 13, 14, 15, 16, 17, 18, 19, 20, 25 solar mass models (default physics and resolution)
- 19 solar mass model: low, medium, high angular resolution; with and without Horowitz correction; monopole versus multipole
- Default resolution: 678 x 128 x 256; 12 energy groups; dendritic grid

(~50 2D models performed: 678 x 128)



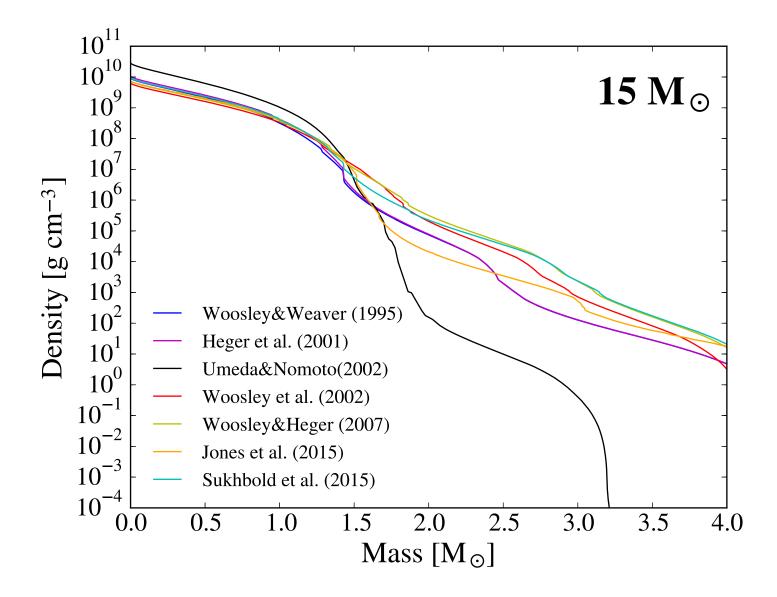


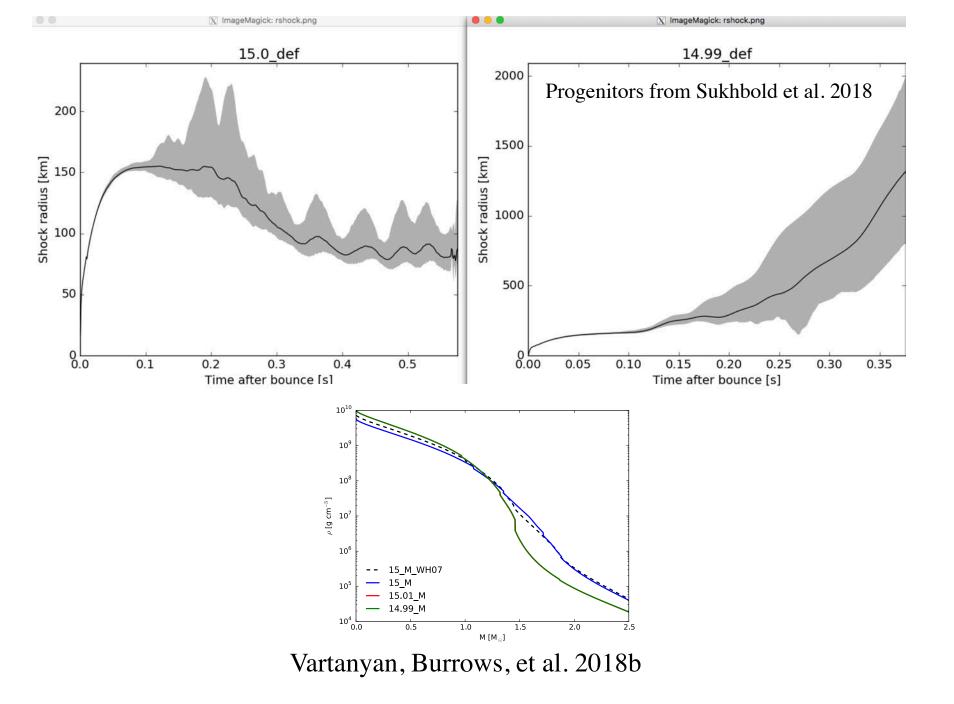


## Important Roles of Progenitor Models:

## Density Structures, Rotational Profiles, Seed Perturbations

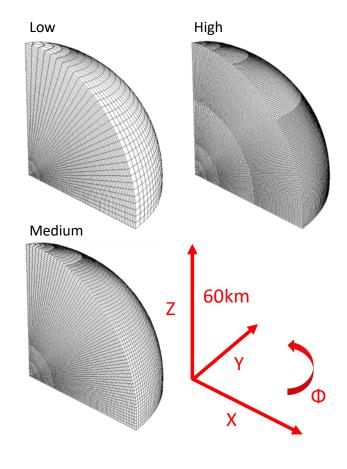
### Different Groups, Same ZAMS Mass

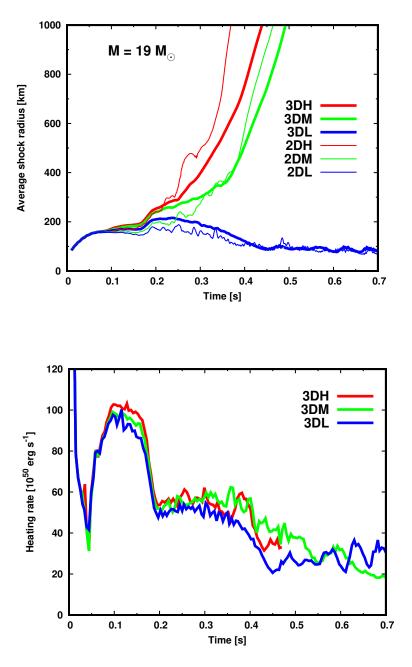




## **Spatial Resolution Dependence**

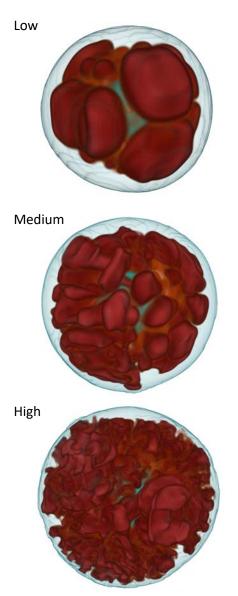
Nagakura et al. 2019

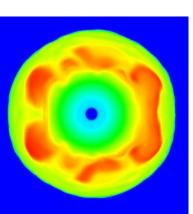


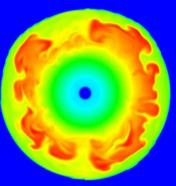


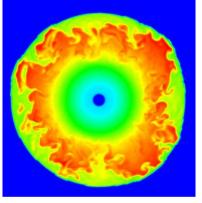
#### 100 ms

#### 200 ms

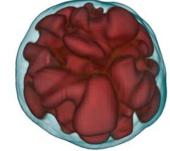


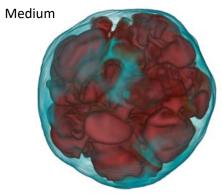




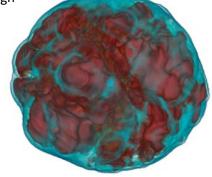


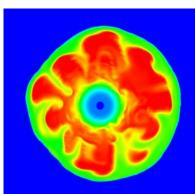
Low

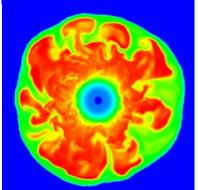


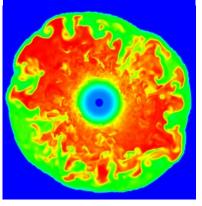


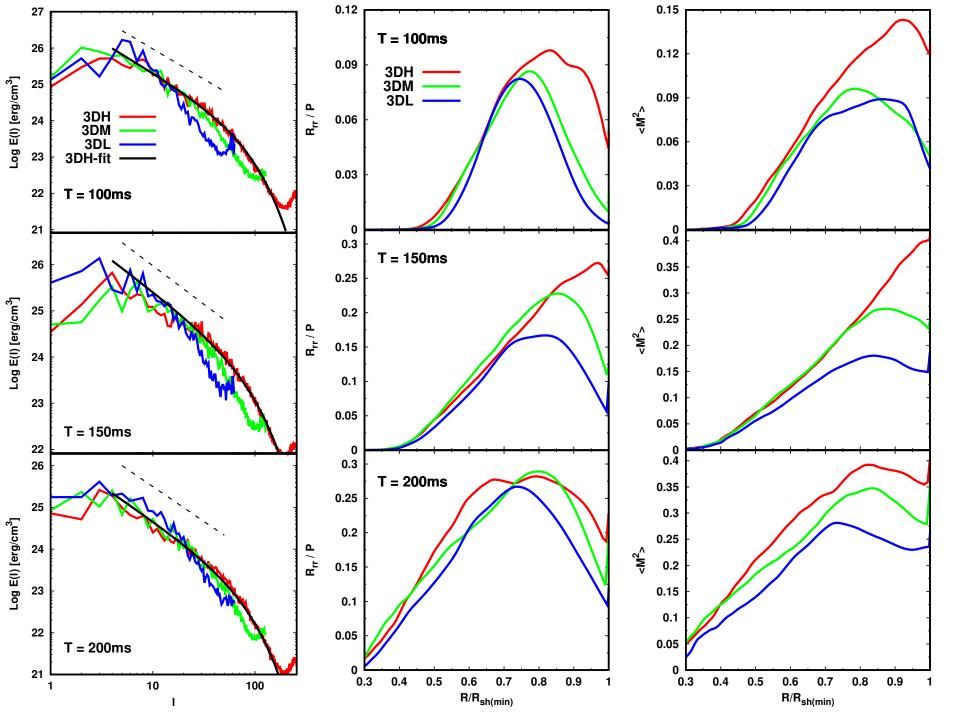
High





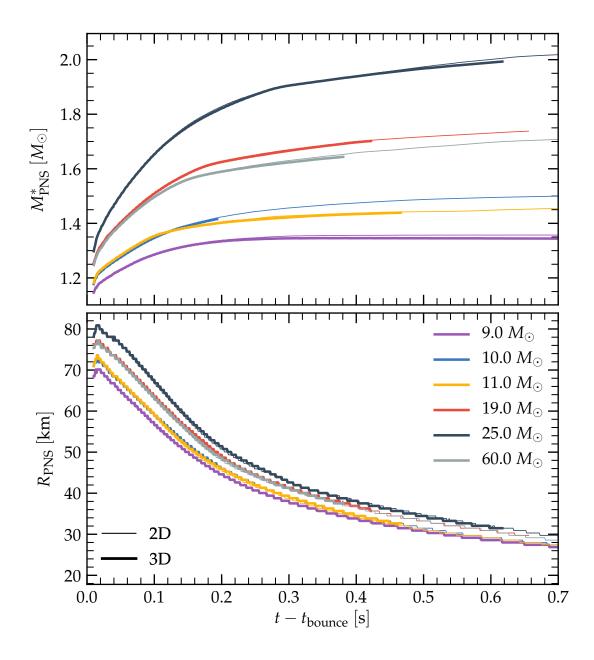






# **New Fornax 3D Simulations**

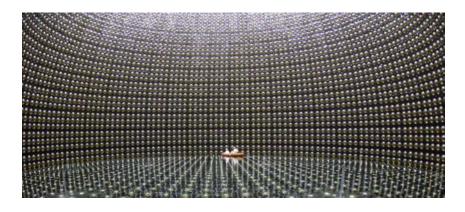
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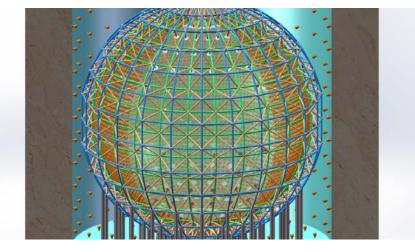
## Supernova Neutrino Detection

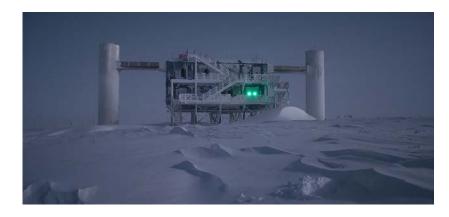
SUPERK, HYPERK, DUNE, JUNO, ICE CUBE

### **SN Neutrino Observatories**

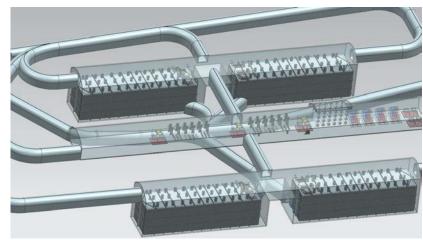


#### Super-Kamiokande (Water Cherenkov)



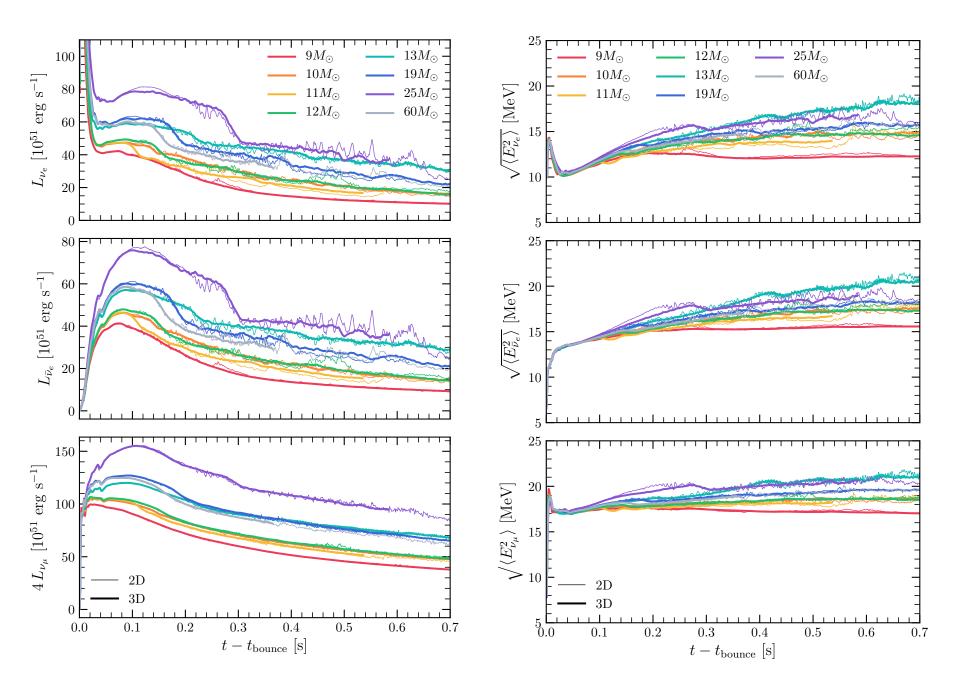


ICECUBE (Longstring Ice)



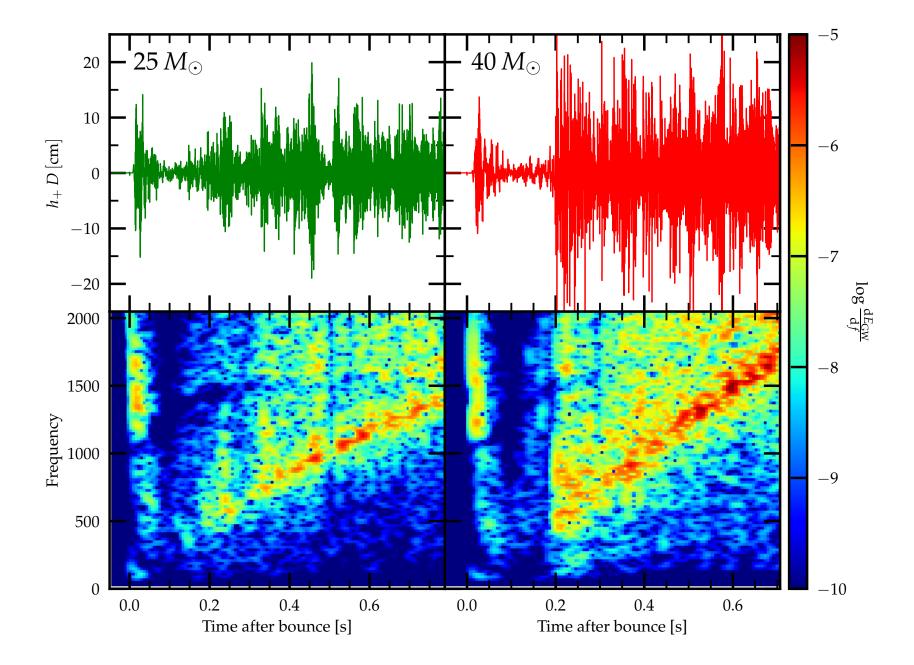
#### JUNO (Hydrocarbon Scintillator)

DUNE (Liquid Argon TPC)



# Gravitational Radiation Signals from Core-Collapse Supernovae

Radice, Morozova, Burrows, Vartanyan et al. 2018-2019



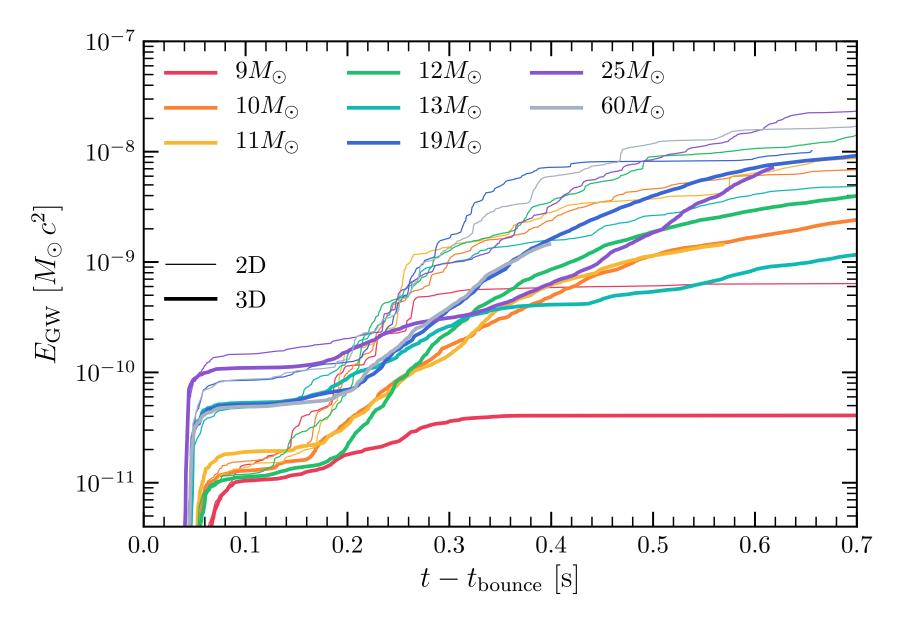




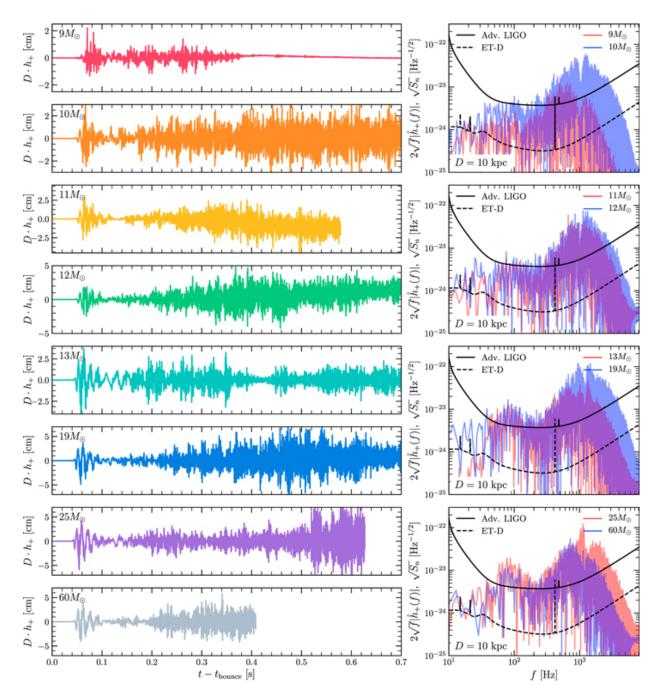




3D (thick) and 2D (thin) Models

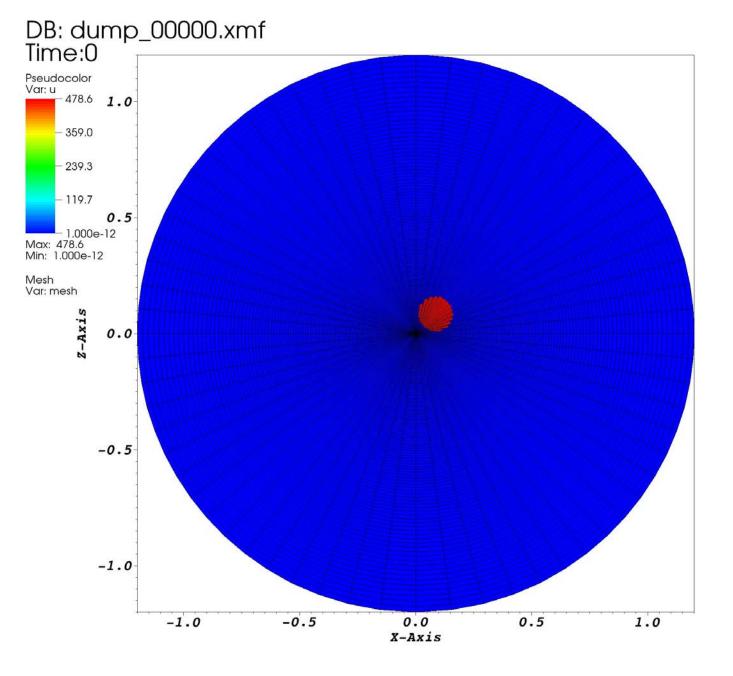


#### Radice et al. 2019



### Core-Collapse Theory: A Status Summary

- Can now perform many 3D simulations per year on HPC resources such as Blue
- Waters!
- Proximity to critical explosion curve amplifies effects of sub-dominant processes, etc.
- Can explain current differences between groups (!?)
- Turbulent convection is Key Enabler of explosion for (almost) all viable mechanisms; turbulent stress, simultaneous accretion and explosion
- Neutrino-driven convection > SASI (when object explodes to yield SN)
- SASI is not a mechanism can't generate much entropy; failed models show SASI (spiral modes)
- Accretion of the Si/O interface
- 3D different from 2D (turbulent pressure, spectrum; scales)!
- Various heating processes (in-medium/many-body, inelastic on electrons, inelastic on nucleons) add "non-linearly"
- Structure factor/many-body corrections! Neutrino-matter interactions!
- **Proto-neutron Star (PNS) Convection** boosts  $v_{\mu}$  neutrino luminosity
- Seed Perturbations
- **Progenitor profiles/structure important!** (e.g., Meakin & Arnett; Couch et al. 2015; B. Muller et al. 2016); Seed Perturbations, Density profiles, Si/O shelfs?
- Rotation!?
- Crucial role for microphysics many-body/structure-factor corrections, inelastic scattering; when near critical curve, small effects are amplified - (partial) origin of differences between groups



user: dradice Thu Dec 20 12:52:41 2018

#### 16 solar mass

### Time = 0.677 s

